

Lecture # 25

Poisson Equation with
magnetic field

Liouville equation

Statement of conservation of probability of $f(\tilde{r}, \tilde{p}, t) d^3\tilde{r} d^3\tilde{p}$

$$\frac{\partial f(\tilde{r}, \tilde{p}, t)}{\partial t} + \tilde{r} \cdot \frac{\partial f}{\partial \tilde{p}} + \tilde{p} \cdot \frac{\partial f}{\partial \tilde{r}} = 0$$

$$+ \frac{\partial f}{\partial \tilde{p}}(\tilde{p}) = 0$$

Since $\mathcal{V}(\tilde{r}, \tilde{p})$ is conserved

[i.e. 6-D (Jacobian) in motion of particles]

what flows into a fixed phase

space volume, must lead to ~~an~~ corresponding increase in particle number (like counting charge in 3-D to get continuity equation)

There is another physical meaning of Liouville equation. The value of

$f(\tilde{r}, \tilde{p}, t)$ along the particle trajectory

does not change.

$$0 = \frac{d}{dt} f(\tilde{r}, \tilde{p}, t) = \left[\tilde{r} \cdot \frac{\partial f}{\partial \tilde{p}} + \tilde{p} \cdot \frac{\partial f}{\partial \tilde{r}} + \frac{\partial f}{\partial t} \right] f(\tilde{r}, \tilde{p}, t) = 0$$

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$$0 = \frac{d}{dt} \int f(\vec{r}, \vec{p}, t) d^3r = \int \left[\frac{\partial f}{\partial t} + \vec{v} \cdot \nabla f + \dot{\vec{p}} \cdot \nabla_p f \right] d^3r =$$

0 - this term vanishes

$$0 = \int \left[\frac{\partial f}{\partial t} + \nabla_p \cdot \left(\vec{p} f \right) - \nabla \cdot \left(\vec{v} f \right) \right] d^3r$$

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equivalent to continuity equation along trajectory

$$0 = \int \left[\frac{\partial f}{\partial t} + \nabla_p \cdot \left(\vec{p} f \right) - \nabla \cdot \left(\vec{v} f \right) \right] d^3r$$

Then conservation form from flux

$$\dot{\vec{r}} = \nabla_p \mathcal{H}(\vec{r}, \vec{p}, t)$$

$$\dot{\vec{p}} = -\nabla_{\vec{r}} \mathcal{H}(\vec{r}, \vec{p}, t)$$

The two forms of the virial theorem are identical by dynamics

$$0 = \left[\tilde{V} \cdot \frac{\partial}{\partial \tilde{r}} + \tilde{p} \cdot \frac{\partial}{\partial \tilde{p}} \right] f(c; \tilde{r}, \tilde{p})$$

Vlasov equation

of motion, satisfy stationary

that are functions of the constants

Thus, a distribution function

where c_i are constants of motion of a particle for Hamilton H_c

$$[H, f(c_i)] = 0$$

dynamics $c_i = c_i(\tilde{r}, \tilde{p})$

Fundamental theorem in Hamiltonian

Poisson Bracket

$$= [H, f] = 0$$

$$= \frac{\partial c}{\partial H} \cdot \frac{\partial c}{\partial \tilde{r}} - \frac{\partial c}{\partial \tilde{r}} \cdot \frac{\partial c}{\partial \tilde{p}} = 0$$

$$0 = \tilde{r} \cdot \frac{\partial c}{\partial \tilde{r}} + \tilde{p} \cdot \frac{\partial c}{\partial \tilde{p}}$$

Stationary Vlasov Equation

Let us consider electro-static

perturbations of a particle in a uniform magnetic field and

Let us take $\vec{B} \parallel \hat{z}$ and

plasma independent of y

$$\vec{B} = B_0 \hat{z}, \quad \vec{A} = X B_0 \hat{y}$$

$$P_y = m v_y + e A_y = \text{constant of motion}$$

$$= m(v_y + e B_0 X)$$

or equate lengthly $X_g = X + Y/c$

(for uniform field $P_y = m \omega_c X_g$)

Other constants of motion are

$$(a) \quad \frac{m}{2} \dot{r}^2 = \frac{1}{2} (v_x^2 + v_y^2) \rightarrow E \text{ (energy/mass)}$$

$$(b) \quad \frac{m}{2} \dot{\theta}^2 \rightarrow \mu_B \equiv \frac{e \hbar}{2m} \text{ (if } B \text{ non-uniform) in space}$$

μ_B is only an adiabatic

invariant, can lead to some modifications which may be treated later

(7)

For now we treat V_2 as a

constant of motion in an unperturbed magnetic field: Also $V_{||} = \pm \sqrt{E - V_2^2}$ constant of motion in this case.

We will consider linear-Vlasov equation

$$f(\vec{r}, \vec{p}, t)$$

$$\left[\frac{\partial}{\partial t} + \vec{v} \cdot \frac{\partial}{\partial \vec{r}} + \frac{e}{c} \vec{v} \times \vec{B} \cdot \frac{\partial}{\partial \vec{p}} \right] f = e \nabla \phi \cdot \frac{\partial f(\vec{r}, \vec{p}, t)}{\partial \vec{p}}$$

$$\nabla^2 \phi = -4\pi e \int d^3 p \sum_i f_i - \int d^3 p \sum_i f_i e$$

$$F = F(x + \frac{\omega_c}{V_{||}}, V_{||}, V_2) = F(x_g, V_{||}, V_2)$$

First just consider single species (say electrons)

$$\frac{\partial}{\partial t} \left| \frac{\partial \tilde{f}}{\partial V} \right|_{\vec{r}} = \frac{\partial}{\partial V} \left| \frac{\partial \tilde{f}}{\partial x_g} \right|_{\vec{r}} + \frac{\partial}{\partial V} \left| \frac{\partial \tilde{f}}{\partial V_{||}} \right|_{\vec{r}} + 2V_2 \frac{\partial \tilde{f}}{\partial V_2}, \text{ and } \therefore$$

Notice:

$$\nabla \phi \cdot \frac{\partial}{\partial \vec{F}} = \frac{\partial \phi}{\partial x} \frac{\partial}{\partial V} + \frac{\partial \phi}{\partial y} \frac{\partial}{\partial V} + \frac{\partial \phi}{\partial z} \frac{\partial}{\partial V} + 2V_2 \frac{\partial \phi}{\partial V_2}$$

(5) by derivative

$$\underline{n} \tilde{v} = c \tilde{b} \times \tilde{\Delta} \tilde{p} \equiv \text{diamagnetic flow}$$

with $\tilde{p} = m \int d^3 p \tilde{v}^2 F = m \int d^3 p (\tilde{v}_x^2 + \tilde{v}_y^2) F$

$$\underline{v} = \tilde{b} \times \tilde{\Delta} \int d^3 p \tilde{v}^2 F = \frac{eB}{eB} \tilde{b} \times \tilde{\Delta} \tilde{p}$$

perpendicular pressure

$$2\pi \tilde{b} \times \tilde{v} \cdot \tilde{\Delta} = 2\pi \tilde{b} \times \left[\tilde{v}_x^2 + (\tilde{v}_y^2 - \tilde{v}_z^2) \right] \cdot \tilde{\Delta}$$

$$d^3 v = d v_x d v_y d v_z$$

$$\tilde{v} = v_x \tilde{b} + v_y \cos \kappa \tilde{y} + v_z \cos \kappa \tilde{z}$$

$$\int d^3 p \tilde{v} \cdot \tilde{\Delta} F(\tilde{v}_x, \tilde{v}_y, \tilde{v}_z) = \int d^3 p \tilde{v} \cdot \tilde{\Delta} F(\tilde{v}_x, \tilde{v}_y, \tilde{v}_z)$$

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(7)

$$\tilde{y}_1 = k_1 \cos \theta \tilde{x} + k_2 \sin \theta \tilde{y} + k_3$$

$$\phi = \phi(t) \exp[ik_x \tilde{x} + ik_y \tilde{y} + ik_z z + i\omega t] \quad \text{c.c.}$$

Take

Let $f = e^{\frac{m}{\rho} \frac{\partial f}{\partial t}} \phi(\tilde{r}, t) + g(\tilde{r}, \lambda)$

note: $\tilde{y} \cdot \nabla \phi = \left[\frac{\partial}{\partial t} - \frac{\partial}{\partial z} - v_{||} \frac{\partial}{\partial z} \right] \phi$

$= e^{\frac{m}{\rho} \left[\frac{\partial}{\partial t} + \frac{\partial}{\partial z} + 2 \frac{\partial}{\partial t} \tilde{y} \cdot \nabla \phi(\tilde{r}, t) \right]}$ Introduce Trick to ease integration

$$\left[\frac{\partial}{\partial t} + \tilde{y} \cdot \frac{\partial}{\partial \tilde{r}} + \frac{\partial}{\partial z} \tilde{y} \cdot \frac{\partial}{\partial \tilde{y}} \right] f$$

$Df/Dt \equiv$ Viscous equation: $(v_{||} \equiv v_z)$ in this case

Now let us examine perturbed

Diagnetic Current

$$j = \sum n e_j \tilde{y} = c b \times \nabla P_{\text{total}}$$

Pressure MHD of fluid theory

$$j = c b \times \nabla P \leftarrow \text{pressure}$$

$j = \sum_j n_j e_j$ $j =$ all species eg. electrons + protons

$$\tilde{k} \cdot \tilde{p} \times \tilde{r}(t) = k_1 v_1 \sin(\omega t + \theta) = k_1 v_1 \sin(\omega t + \theta)$$

$$\tilde{k} = k_{||} \tilde{z} + k_{\perp} \cos \theta \tilde{x} + k_{\perp} \sin \theta \tilde{y}$$

For $\phi(t', r') = \phi_k \exp[-i\omega t' + i\tilde{k} \cdot \tilde{r}']$

~~$$R_g(z) = R_g(t) + v_1 z$$~~

$$\tilde{V} = v_1 \cos \psi \tilde{x} + v_1 \sin \psi \tilde{y}$$

$$\tilde{V}(z) = v_1 \cos(\psi - \omega z) \tilde{x} + v_1 \sin(\psi - \omega z) \tilde{y}$$

$$\tilde{r}(z) = R_g(z) + \tilde{p} \times \tilde{V}(z) - \tilde{p} \times \tilde{V} / \omega_c + \tilde{r} + v_1 z \tilde{z}$$

$$r(z) = R_g(z) + \tilde{p} \times \tilde{V}(z) + \tilde{p} \times \tilde{V} / \omega_c + v_1 z \tilde{z}$$

$$\tilde{r}(0) = \tilde{r}, \quad t' - z = z$$

where $\tilde{r}(0) = \tilde{r}$

$$\int dt' \phi(t', r(t'-z), \tilde{V}(t'-z))$$

$$g(r, v, t) = ie \left[\frac{k_y}{m} \frac{\partial F}{\partial v_y} + k_z \frac{\partial F}{\partial v_z} + (\omega - k_z v_z) \frac{\partial F}{\partial v_z} \right]$$

Integrate by method of characteristics:

$$F = F(x_y, v_x, v_y, v_z)$$

Dt

$$Dg(r, v, t) = ie \phi \left[\frac{k_y}{m} \frac{\partial F}{\partial x_y} + k_z \frac{\partial F}{\partial v_z} + (\omega - k_z v_z) \frac{\partial F}{\partial v_z} \right]$$

functions of constants of motion in equilibrium orbits

Then $\tilde{\nabla}_r \phi \rightarrow i\tilde{k} \phi_k$

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$$\sum_{n=0}^{\infty} I_n(x) = \exp[x \sin \theta]$$

note

$$\phi_k = \int_0^{\infty} e^{-i\omega t} dt \exp[-i(\omega - k)v] \exp[i\sin(\omega + \theta) t]$$

$$\phi_k = \int_0^{\infty} e^{-i\omega t} dt \exp[-i\omega t + i k v t + i \sin(\omega + \theta) t]$$

$$= \int_0^{\infty} e^{i t [-\omega + k v + \sin(\omega + \theta)]} dt$$

Thus we have

Then

$$\phi(x) = \int_0^{\infty} h(x) dx \exp[i k(x) - x]$$

where $\frac{d\phi(x)}{dx} < 1$, also $\frac{d^2\phi}{dx^2} < 1$

We have assumed in WKB approximation that $\phi(x) = \exp(i \int h(x) dx)$

essentially have assumed in WKB approximation that $\phi(x) = \exp(i \int h(x) dx)$

taken a