

Lecture # 19

Ballooning Mode Formalism

Pressure Driven Modes

At zero beta only a flute can be unstable, and if

$\frac{d}{d\ell} \int \frac{d\ell}{B} < 0$, stability is guaranteed.

However, a "ballooning" mode, when field line bends in bad curvature region, instability can arise if there is sufficient pressure gradient to drive the mode.

Basically, stabilization arises from bending energy, while destabilization arise from pressure drive where there is "bad"

curvature $\kappa \frac{\partial p}{\partial \ell} > 0$.

First let us look at a cylinder and derive a sufficient \mathcal{Q}

condition for instability (Suydam condition)

consider the terms

\underline{Q}_\perp and $-2(\underline{S}_\perp \cdot \nabla \rho) \underline{\chi} \cdot \underline{\rho}$ in energy principle

$$\delta W_I \approx \int d^3r \left[\frac{|\underline{Q}_\perp|^2}{2} - (\underline{S}_\perp \cdot \nabla \rho) (\underline{\chi} \cdot \underline{\rho}^*) \right]$$

Let: $\underline{Q}_\perp = \nabla \times (A_\parallel \underline{b}) \approx -\underline{b} \times \nabla A_\parallel$

$$\frac{\partial \underline{S}_\perp}{\partial \underline{B}} = \frac{\underline{b} \times \nabla \phi}{B} = \gamma \underline{S}_\perp$$

$$\underline{b} \cdot (\underline{E} + \underline{v} \times \underline{B}) = 0 = -\frac{\partial A_\parallel}{\partial t} - \frac{\partial \phi}{\partial s}, \quad ds = \begin{array}{l} \text{distance} \\ \text{along} \\ \text{field} \\ \text{line} \end{array}$$

$$A_\parallel = \frac{-1}{\gamma} \frac{\partial \phi}{\partial s} \quad \text{--- } \gamma A_\parallel$$

$$\therefore \underline{Q}_\perp = \frac{\underline{b} \times \nabla}{\gamma} \frac{\partial \phi}{\partial s}$$

$$\underline{S}_\perp = \frac{\underline{b} \times \nabla}{\gamma B} \phi$$

(set $\gamma=1$ as common factor of all terms)

$$\delta W_I \approx \frac{1}{2} \int d^3r \left[\frac{|\underline{b} \times \nabla \frac{\partial \phi}{\partial s}|^2}{B^2} - \frac{2(\underline{b} \times \nabla \phi \cdot \nabla \rho) (\underline{\chi} \cdot \underline{b} \times \nabla \phi)^*}{B^2} \right]$$

Near resonance we look for solution

where $\frac{\omega}{\omega r} \gg \frac{m}{r}$, $\underline{\chi} = -\frac{B_\theta}{B^2} \frac{1}{r} \underline{r}$
 $B_z \gg B_\theta$

$$\frac{\partial \phi}{\partial s} = \frac{B_0}{rB} (m - nq(r))$$

resonance condition $m - nq(r_s) = 0$

Euler - Lagrange equation

$$\frac{\partial}{\partial r} \frac{B_0^2(r)m^2}{r^2 B^2} \left(1 - \frac{n}{m} q(r)\right)^2 \frac{\partial \phi}{\partial r} - \frac{2m^2}{r^3} \frac{\partial p}{\partial r} \frac{B_0^2}{B^2} \phi = 0$$

near resonance surface

$$\frac{m}{n} = q(r_s), \quad r = r_s + \delta r$$

$$\left(1 - \frac{n}{m} q(r)\right) \approx \delta r q'(r_s) / q(r_s)$$

$$\frac{B_0^2(r_s)}{r_s^2} \frac{q'(r_s)^2}{q^2(r_s)} \frac{\partial}{\partial r} \delta r^2 \frac{\partial \phi}{\partial r} - \frac{2}{r_s^3} \frac{\partial p(r_s)}{\partial r} \frac{B_0^2(r_s)}{B^2(r_s)} \phi = 0$$

search for a solution of

equation $\frac{\partial}{\partial r} \delta r^2 \frac{\partial \phi}{\partial r} + D \phi = 0$; $D = \frac{2}{r_s} \frac{\partial p(r_s)}{\partial r} \frac{q^2(r_s)}{B^2 \left(\frac{\partial q(r_s)}{\partial r}\right)^2}$

$\phi = r^\nu$, upon substitution

$$\nu(\nu+1) + D = 0; \quad \nu = -\frac{1}{2} \pm \left[\frac{1}{4} - D\right]^{1/2}$$

Complex, oscillatory solution if $D > \frac{1}{4}$
as $\delta r \rightarrow 0$

$$\delta r^{-1/2 \pm i[D-1/4]^{1/2}} = \frac{\exp[\pm i(D-1/4)^{1/2} \ln|\delta r|]}{\delta r^{1/2}}$$

↑
oscillatory as $\delta r \rightarrow 0$

Therefore instability guaranteed if

$$4D = \frac{\gamma}{\Gamma_s} \frac{\omega_p^2}{\omega^2} > 1$$

Saydam criterion

Recall Newcomb's theorem that notes if Euler-Lagrange Equation of SW is oscillatory, system is unstable

(4) If Newco Suydam criterion gives stability ($D < 1/4$), instability still possible from non-local source. However Euler equation has to be integrated over entire radius (or at least within region of ~~with~~ rational surface)

Ballooning Mode Equation

(1) Leads to a more general sufficient instability condition for a tokamak with shifted magnetic surfaces (Mercier condition)

$$4D_M > 1, \quad D_M = - \frac{8r_s}{B_0^2(r_s)} \frac{\partial p(r_s)}{\partial r} \frac{1 - q^2(r_s)}{q^2(r_s)}$$

The extra q^2 is indicative that a magnetic well forms if

$q^2 > 1$. For $q < 1$, (no magnetic well, The extra q^2 factor

difficult to calculate simply (not here) (5)

However, we will outline the ballooning formalism that has been found to be very informative for instability analysis in confined systems:

The SW energy integral has the following terms (when magnetic compressional terms are removed)

$$\delta W = \frac{1}{2} \int d^3r \left[|\tilde{Q}_\perp|^2 - 2(\tilde{S}_\perp \cdot \nabla \rho)(\tilde{S}_\perp^* \cdot \tilde{x}) \right]$$

For short wavelength, let us represent these perturbations in the form (as before)

$$\tilde{S}_\perp = \frac{\tilde{b} \times \nabla \phi}{\gamma B}, \quad \tilde{Q}_\perp = \nabla \times \tilde{b} A_\parallel \approx -\tilde{b} \times \nabla A_\parallel$$

ideal Ohm's law. $\tilde{E}_\parallel = -\frac{\partial A_\parallel}{\partial t} - \nabla \phi \cdot \tilde{b} = 0$

$$\therefore A_\parallel = \frac{-\tilde{b} \cdot \nabla \phi}{\gamma}$$

$$\therefore \tilde{S}_\perp' \equiv \gamma \tilde{S}_\perp = \tilde{b} \times \nabla \phi / B$$

$$\tilde{Q}_\perp' = \gamma \tilde{Q}_\perp = \tilde{b} \times \nabla \frac{\partial \phi}{\partial s}$$

ds : distance along field line

(now γ can be set to unity without loss of generality) (5)

$$\delta W = \frac{1}{2} \int d^3r \left(\frac{\partial \mathcal{L}}{\partial \beta} \right)^2 \left[\begin{array}{l} \left(\vec{b} \times \vec{\nabla} \beta - g(\psi) \vec{b} \times \vec{\nabla} \theta - \vec{b} \times \vec{\nabla} \psi g'(\theta) \right) \left| \frac{\partial \phi}{\partial \beta} \right|^2 \\ \cdot \left(\vec{b} \times \vec{\nabla} \beta - g(\psi) \vec{b} \times \vec{\nabla} \theta - \vec{b} \times \vec{\nabla} \psi g'(\theta) \right) \left| \frac{\partial \phi}{\partial \beta} \right|^2 \\ - 2 \left| \frac{\partial \phi}{\partial \beta} \right|^2 \left[\chi_\psi + \frac{\chi_g}{B} \left(\vec{\nabla} \psi \cdot (\vec{\nabla} \beta + g(\psi) \vec{\nabla} \theta) + g'(\theta) |\vec{\nabla} \psi|^2 \right) \right] \end{array} \right]$$

Aside:

It is interesting that parallel current drive (kink) does not appear in this limit.

$$\begin{aligned} J_{\parallel} (\beta_{\perp}^* \times b) \cdot Q_{\perp} &\simeq J_{\parallel} (\vec{b} \times \vec{\nabla} \phi) \times \vec{b} \cdot \vec{b} \times \vec{\nabla} \frac{\partial \phi}{\partial \beta} \\ &= J_{\parallel} \vec{\nabla}_{\perp} \phi \cdot (\vec{b} \times \vec{\nabla}_{\perp} \phi) \frac{\partial \phi}{\partial \beta} \\ &= J_{\parallel} \left(\frac{\partial \beta}{\partial \beta} \right)^2 \left[\vec{\nabla}_{\perp} \beta \cdot \vec{b} \times \vec{\nabla}_{\perp} \beta \right] \frac{\partial \phi}{\partial \beta} = 0 \end{aligned}$$

The Euler-Lagrange Equation

The Euler-Lagrange Equation is $[\vec{\nabla} \beta \cdot \vec{\nabla} \theta = 0]$

$$\frac{\partial}{\partial \beta} \left[\vec{\nabla}_{\perp} \beta \cdot \vec{\nabla}_{\perp} \beta + g^2(\psi) \vec{\nabla}_{\perp} \theta \cdot \vec{\nabla}_{\perp} \theta + (|\vec{\nabla} \psi|^2 g'^2 \theta^2 + 2g(\psi) \vec{\nabla} \theta \cdot \vec{\nabla} \psi) \frac{\partial \phi}{\partial \beta} \right] \frac{\partial \phi}{\partial \beta}$$

$$+ 2 \left[\frac{\partial \mathcal{L}}{\partial \psi} \left[\chi_\psi + \frac{\chi_g}{B} (g(\psi) \vec{\nabla} \psi \cdot \vec{\nabla} \theta + g'(\theta) |\vec{\nabla} \psi|^2) \right] \right] \phi = 0$$

(more generally $\theta \rightarrow \theta - \theta_0$)