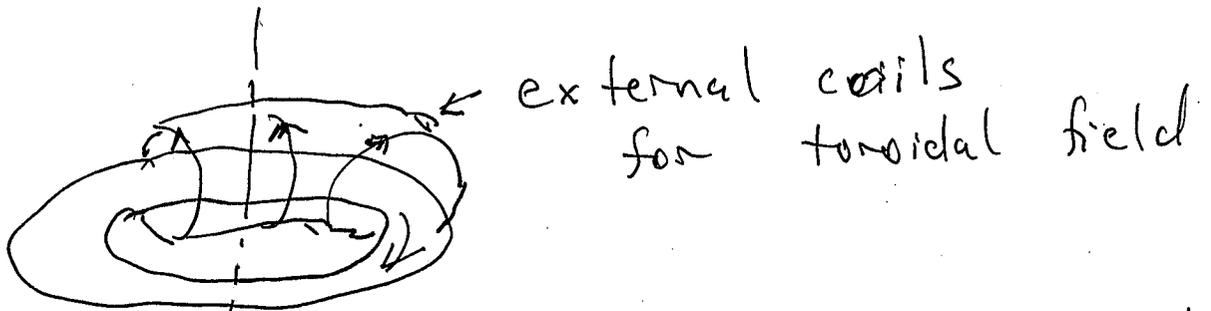


# Lecture # 10.

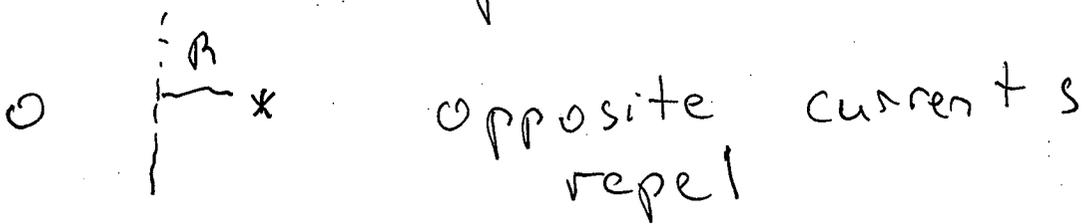
Toroidal Grad-Shafranov

Equation

### 3-D confinement, with symmetry



Note that a toroidal current exhibits a "hoop" stress



$$\text{Force} \approx I B L, \quad L = 2\pi R$$

$$B \sim I/R$$

$$\text{Force} \approx I^2 \quad \text{cause expansion}$$

A vertical field,  $B_y$ , with a counter acting force needs to be supplied:

$$B_y \approx \frac{a}{R} \quad B_p \approx \frac{I}{R}$$

We represent the magnetic field as:

$$\vec{B} = \vec{\nabla} \psi \times \vec{\nabla} \psi - I \vec{\nabla} \psi$$

(mixed representation)  
Poloidal Field  $\equiv$  contra variant vector

Toroidal Field  $\equiv$  co-variant vector

Pressure  $\equiv$  only a function of  $\psi$

$$\vec{\nabla} P = \vec{J} \times \vec{B}$$

$$\therefore \vec{B} \cdot \vec{\nabla} P = 0$$

As we move along field line, which generates a nested flux tube, pressure cannot vary. Hence

$$P = P(\psi), \quad \vec{\nabla} P = \vec{\nabla} \psi \frac{\partial P}{\partial \psi}$$

From  $\nabla \cdot \vec{J} = 0$ , and toroidal symmetry, and  $\vec{J} \cdot \nabla \psi = 0$  (as  $\nabla \psi \frac{\partial \rho}{\partial \psi} = \vec{J} \times \vec{B}$ ) it follows that the poloidal current is parallel to the poloidal magnetic field.

$$\vec{J}_p = \alpha \nabla \rho \times \nabla \psi \quad \left( \begin{array}{l} \text{only way } \vec{J}_p \perp \\ \text{to both } \nabla \rho \text{ and } \nabla \psi \end{array} \right)$$

Further, from ampere's law we take  $\int \vec{B} \cdot d\vec{l} = \int \vec{J} \cdot d\vec{s} = I_{\text{enc}} 2\pi r$  in toroidal symmetry direction, and it must be equal to the total poloidal current,  $I(\psi)$  through the ribbon area ~~slice~~ element generated by the element

$$dA = \frac{\nabla \rho \cdot \nabla \psi}{J} \Rightarrow 2\pi \frac{\nabla \theta \cdot \nabla \psi}{J}$$

$$\therefore -2\pi I = I(\psi) = \int \vec{J}_p \cdot dA = 2\pi \int \alpha \frac{\nabla \theta \cdot \nabla \psi}{J} d\psi = 2\pi \int \alpha d\psi$$

$$\therefore \alpha = \frac{-I(\psi)}{J}$$

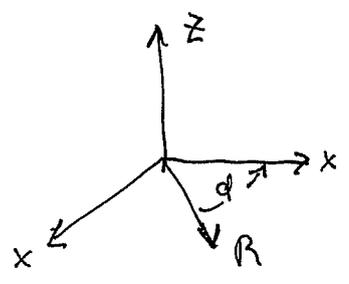
The current  $I(\psi)$  includes the poloidal current in the plasma as well as the poloidal current in the magnets forming the vacuum toroidal field.



$$\nabla \cdot \left( \frac{\nabla \psi}{R^2} \right) = R^2 \frac{\partial p}{\partial \psi} + I(\psi) \frac{\partial I(\psi)}{\partial \psi}$$

An example of a solution is the Solov'ev solution, when  $I(\psi) = 0$ . In this case there is no external magnets supplying toroidal fields. He also chooses  $\frac{\partial p}{\partial \psi} = \frac{\partial p}{\partial \psi} \Big|_0 = \text{constant}$

In cylindrical geometry



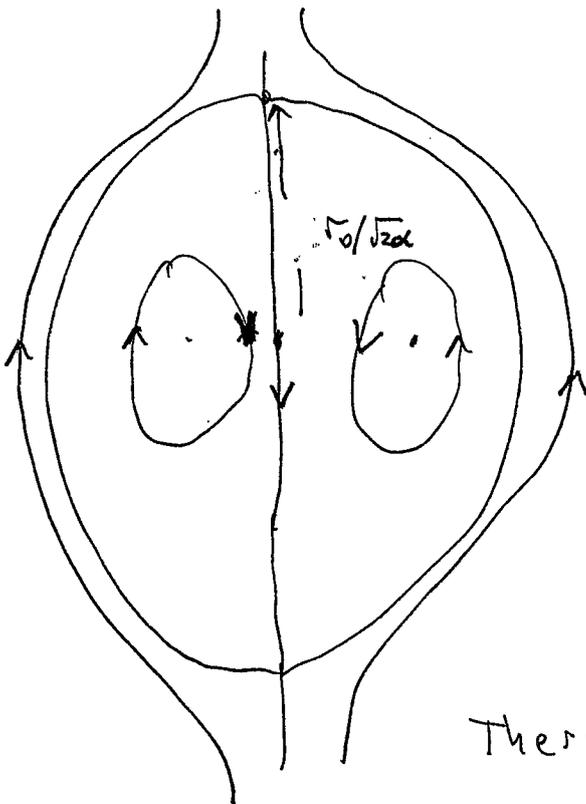
$(R, \phi, z)$   
the <sup>G-S</sup> equation takes the form

$$-R \frac{\partial}{\partial R} \frac{1}{R} \frac{\partial \psi}{\partial R} + \frac{\partial^2 \psi}{\partial z^2} = - \frac{\partial p}{\partial \psi} \Big|_0$$

The A solution is of the form:

$$\psi = \psi_0 \frac{r^2}{r_0^4} (2r_0^2 - r^2 - 4d^2 z^2)$$

(determine how these parameters depend on  $\frac{\partial p}{\partial \psi_0}$ )



This solution describes plasma on closed field lines ( $\psi > 0$ ) and open field lines ( $\psi < 0$ )

There are magnetic nulls

at  $r=0, z = \pm r_0/d\sqrt{2}$ . Also at O-point, ( $z=0, r=r_0$ )

Typically we are interested in case where  $p=0$  on separatrix ( $\psi=0$  in this case) and  $p=0$  on open field lines

For this case the solution of the problem is incomplete, in that the required vertical needs to be determined. In general this is a hard problem.

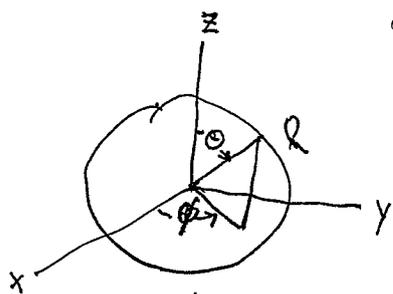
The one straight-forward case is when the separatrix is a sphere ( $\alpha = 1/2$ ). The matching analytic solution can be found. It is easiest to show in spherical coordinates ( $\rho^2 = r^2 + z^2$ ) where G-S equation is

$$\frac{\partial^2 \psi}{\partial \rho^2} + \frac{1}{\rho^2 \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial \psi}{\partial \theta} \right) = -\frac{2}{\rho^2 \sin^2 \theta} \frac{\partial \psi}{\partial \phi} - \frac{I(\psi) I(\psi)}{\rho^2}$$

If rhs vanishes;

$$\psi = \left( \frac{A}{\rho} + B \rho^2 \right) \cos \theta$$

for  $\rho > \sqrt{2} r_0$



This solution can be matched to the interior solution, (an exercise you need to do for your homework). Note  $B \rho^2 \cos \theta$  is the flux function for a uniform vertical field.