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of the Vlasov-Maxwell System**

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# Exact Nonlinear Travelling Solutions of the Vlasov-Maxwell System

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Exact time dependent solutions of a system of charged particles with their self-consistent electromagnetic fields are obtained by solving the fully nonlinear Vlasov-Maxwell set of equations.

Exact nonlinear solutions to coupled field theories are generally rare. For the Vlasov-Maxwell system, two kinds of time-independent solutions have been reported; 1) the electrostatic solutions of the Bernstein-Green-Kruskal (BGK) type,<sup>1</sup> and 2) the magnetic solution of the type obtained by Pfirsch,<sup>2</sup> Laval-Pellat and Vuillemin,<sup>3</sup> and Marx.<sup>4</sup> In this letter, we present a time-dependent fully electromagnetic solution to the Vlasov-Maxwell equations which describes a system of charged particles interacting with their self-consistent fields.

We consider an electron-ion plasma embedded in a constant external magnetic field  $\mathbf{B}_{\text{ext}} = \hat{e}_z B_0$ . Assuming the field quantities to vary only in the  $x$ -direction ( $\partial/\partial y = 0 = \partial/\partial z$ ), the relevant equations describing the system can be written as

$$\frac{\partial f_e}{\partial t} + v_x \frac{\partial f_e}{\partial x} - \frac{e}{m_e} \left[ \mathbf{E} + \frac{\mathbf{v}}{c} \times (\mathbf{B} + \hat{e}_z B_0) \right] \cdot \frac{\partial f_e}{\partial \mathbf{v}} = 0 \quad (1a)$$

$$\frac{\partial f_i}{\partial t} + v_x \frac{\partial f_i}{\partial x} + \frac{e}{m_i} \left[ \mathbf{E} + \frac{\mathbf{v}}{c} \times (\mathbf{B} + \hat{e}_z B_0) \right] \cdot \frac{\partial f_i}{\partial \mathbf{v}} = 0 \quad (1b)$$

$$\frac{\partial E_x}{\partial x} = 4\pi\rho = -4\pi e \int dv (f_e - f_i) \quad (2)$$

$$\frac{\partial B_x}{\partial x} = 0 \quad (3)$$

$$\frac{\partial}{\partial x}(\hat{e}_x \times \mathbf{E}) + \frac{1}{c} \frac{\partial \mathbf{B}}{\partial t} = 0 \quad (4)$$

$$\frac{\partial}{\partial x}(\hat{e}_z \times \mathbf{B}) - \frac{1}{c} \frac{\partial \mathbf{E}}{\partial t} = \frac{4\pi}{c} \mathbf{J} = -\frac{4\pi e}{c} \int d\mathbf{v} \mathbf{v}(f_e - f_i) \quad (5)$$

where Eqs. (1a) and (1b) are the Vlasov equations for the evolution of the electron and ion distribution functions  $f_e$  and  $f_i$  in the presence of the external field  $B_0$  and the self-consistently generated fields  $\mathbf{E}$  and  $\mathbf{B}$  which obey Maxwell's equations (2-5). In Eqs. (1-5),  $-e$  is the charge of the electron,  $m_e$  and  $m_i$  are the electron and ion masses, and  $\rho$  and  $\mathbf{J}$  are the plasma charge and current densities.

The search for the solution begins by proposing displaced Maxwellian distributions for the particles

$$f_{e,i} = \frac{n_0}{\pi^{3/2} v_{e,i}^3} \exp \left[ -\frac{(\mathbf{v} - \mathbf{u}_{e,i})^2}{v_{e,i}^2} \right] g_{e,i}(x - Ut) \quad (6)$$

where  $n_0$  is a measure of the ambient density,  $v_{e,i}$  and  $\mathbf{u}_{e,i}$  are respectively the thermal speeds and the drift velocities, and the entire space-time dependence is contained in the factor  $g_{e,i}$  which depends only on the variables  $\eta = x - Ut$ . Making use of Eq. (6), one obtains (explicit arguments will be dropped) the expression for the charge and current densities,

$$\rho = -en_0(g_e - g_i) , \quad (7)$$

$$\mathbf{J} = -en_0(\mathbf{u}_e g_e - \mathbf{u}_i g_i) , \quad (8)$$

which implies that the choice

$$g_e = g_i = g \quad , \quad E_x = 0 = B_x \quad (9)$$

trivially satisfies Eqs. (2) and (3). Using Eqs. (6-9) and remembering that  $\partial/\partial t = -Ud/d\eta$ ,  $\partial/\partial x = d/d\eta$ , the remaining nontrivial components of Maxwell's equations take the form

$$\frac{d}{d\eta} \left[ E_y - \frac{U}{c} B_z \right] = 0 \quad (10)$$

$$\frac{d}{d\eta} \left[ E_z - \frac{U}{c} B_y \right] = 0 \quad (11)$$

$$\frac{dB_z}{d\eta} = +\frac{4\pi e}{c}n_0g(u_y^e - u_y^i) + \frac{U}{c}\frac{dE_y}{d\eta} \quad (12)$$

$$\frac{dB_y}{d\eta} = -\frac{4\pi e}{c}n_0g(u_z^e - u_z^i) + \frac{U}{c}\frac{dE_z}{d\eta} \quad (13)$$

Equations (10) and (11) immediately yield

$$E_y = \frac{U}{c}B_z + E_0 \equiv \frac{U}{c}[B_z + B_0] \quad (14)$$

and

$$E_z = -\frac{U}{c}B_y \quad (15)$$

which guides us to the interpretation that  $U$  must be a drift in the  $x$ -direction, and in the moving frame, the effective electric field remains zero. Making use of Eqs. (6), (14) and (15),

$$\mathbf{u}_{e,i} = \hat{e}_z u_z^{e,i} + \hat{e}_y u_y^{e,i} + \hat{e}_x U \quad (16)$$

and some straightforward algebra, the Vlasov equations reduce to [ $T_{e,i} = 0.5 m_{e,i} v_{e,i}^2$  is the temperature]

$$\frac{1}{g_e} \frac{dg_e}{d\eta} = \frac{e}{cT_e} [B_y u_z^e - (B_z + B_0) u_y^e] \quad (17)$$

$$\frac{1}{g_i} \frac{dg_i}{d\eta} = \frac{e}{cT_i} [i B_y u_z^i + (B_z + B_0) u_y^i] \quad (18)$$

which can satisfy the constraint  $g_e = g_i = g$  only if

$$\frac{u_z^e}{T_e} = -\frac{u_z^i}{T_i}, \quad \frac{u_y^e}{T_e} = -\frac{u_y^i}{T_i} \quad (19)$$

leading to the single Vlasov Equation [ $\overline{B}_{y(z)} = (e u_{z(y)}^e / c T_e) B_{y(z)}$ ]

$$\frac{1}{g} \frac{dg}{d\eta} = \overline{B}_y - \overline{B}_z \quad (20)$$

to be solved in conjunction with

$$\frac{d}{d\eta} \overline{B}_z = +\frac{2}{\delta^2} \left( \frac{u_y^e}{u^e} \right)^2 g \quad (21)$$

$$\frac{d}{d\eta} \overline{B}_y = -\frac{2}{\delta^2} \left( \frac{u_z^e}{u^e} \right)^2 g \quad (22)$$

which are Eqs. (12) and (13) rewritten in the context of Eqs. (14), (15) and (19) where

$$\delta^2 = \frac{c^2}{\omega_e^2} \frac{v_e^2}{u^{e^2}(1+\tau)} (1 - U^2/c^2) \quad (23)$$

with  $\omega_e^2 = (4\pi n_0 e^2/m_e)$ ,  $u^{e^2} = u_z^{e^2} + u_y^{e^2}$  and  $\tau = T_i/T_e$  as the ion-electron temperature ratio.

The set of Eqs. (20-22) allows the following self-consistent solutions

$$g = \text{sech}^2 \frac{x - Ut}{\delta}, \quad (24)$$

$$B_y = -\frac{u_z^e}{u^e} \left( \frac{cT_e}{eu_e} \frac{2}{\delta} \right) \tanh \frac{x - Ut}{\delta}, \quad (25)$$

$$B_z = -B_0 + \frac{u_y^e}{u^e} \left( \frac{cT_e}{eu_e} \frac{2}{\delta} \right) \tanh \frac{x - Ut}{\delta}, \quad (26)$$

$$E_y = \frac{U}{c} (B_z + B_0), \quad (27)$$

and

$$E_z = -\frac{U}{c} B_y, \quad (28)$$

which are travelling pulses of width  $\delta$ . The width of the pulse (or the shock front if  $\delta$  is very small) is a strong function of the plasma properties (density, temperature and the drift speeds) and can vary over a large range. For  $B_0 = 0$ , the electromagnetic fields are purely transverse ( $E \cdot B = 0$ ) and become more and more light-like as  $U \rightarrow c$ . However, in this relativistic limit, we must use the relativistic correct form for the Vlasov equation ( $\mathbf{p}$  is the momentum),

$$\frac{\partial f(\mathbf{x}, \mathbf{p}, t)}{\partial t} + \mathbf{v} \cdot \frac{\partial f(\mathbf{x}, \mathbf{p}, t)}{\partial \mathbf{x}} + q \left[ \mathbf{E} + \frac{\mathbf{v} \times \mathbf{B}}{c} \right] \cdot \frac{\partial f}{\partial \mathbf{p}} = 0 \quad (29)$$

with the velocity-momentum relation

$$\mathbf{v} = \frac{\mathbf{p}}{\gamma m} = \frac{\mathbf{p}}{m} \left[ 1 + \frac{p^2}{m^2 c^2} \right]^{-1/2} \quad (30)$$

which introduces some complication in the calculation of the plasma current. This difficulty can be readily overcome by assuming that the flow speed  $U$  can be relativistic while the

temperature of the plasma remains nonrelativistic. In that case  $\gamma$  can be simply replaced by  $\gamma_0 = (1 - \frac{U^2}{c^2})^{-1/2} = (1 + p_U^2/m^2 c^2)^{1/2}$  (where  $p_U = \gamma_0 m U$ ) and the entire analysis remains valid provided the plasma frequency  $\omega_e$  is replaced by the relativistic form  $\omega_{er} = (4\pi n_0 e^2 / \gamma_0 m)^{1/2}$ . It is also convenient to rewrite Eq. (23) as

$$\delta^2 = \frac{1}{\gamma_0^2} \frac{c^2}{\omega_{er}^2} \frac{v_e^2}{u_e^2 (1 + \tau)}. \quad (31)$$

In this paper, we have presented exact travelling pulse-like solutions of the fully nonlinear Vlasov-Maxwell system. Expressions for the self-consistent fields and the particle distribution function are explicitly displayed in Eqs. (24-28). Those solutions should find widespread applications in understanding the phenomena taking place in the laboratory as well as in the astrophysical plasmas; they should also serve as reference solutions for nonlinear numerical calculations.

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