

Simulation of the interaction between Alfvén waves and fast particles in stellarators

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Systems
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Introduction

- The hybrid model

CKA–EUTERPE Hybrid code

- MHD description

- Kinetic description

- Interface between the codes

Benchmarks

- Benchmark 1

- Benchmark with NOVA

- GYGLES benchmark

Stellarator results

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Benchmark with NOVA

GYGLES benchmark

Stellarator results

Alfvén waves

- ▶ Alfvén waves are present in magnetised plasmas
- ▶ There are global eigenmodes in the gaps of the Alfvén wave spectrum
- ▶ These modes could be driven unstable by energetic particles

CKA-EUTERPE Hybrid Code

- ▶ To describe wave particle interaction
- ▶ In real 3D geometry of stellarators
- ▶ Tool for perturbative stability analysis

The hybrid approach

- ▶ The particles are described by kinetic equations
- ▶ The Alfvén waves are calculated from MHD
- ▶ The particles move in the calculated wavefield
- ▶ Nonlinear: the energetic particles modify the wavefield

Why do we need a hybrid model?

- ▶ Kinetic simulation of Alfvén waves is complicated
- ▶ We need a numerical tool to determine stability
- ▶ Better physical model than present hybrid codes (CAS3D-K)

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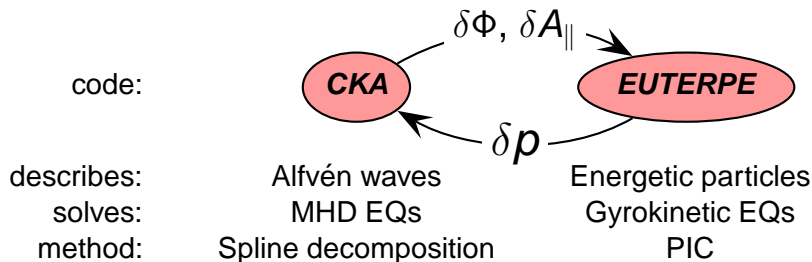
Benchmark 1

Benchmark with NOVA

GYGLES benchmark

Stellarator results

CKA-EUTERPE Hybrid code



- ▶ Real 3D magnetic geometry
- ▶ No assumptions on particle orbits

- ▶ Linearised reduced MHD equations
- ▶ Transformed to an eigenvalue problem:

$$\omega^2 \left[\nabla \cdot \left(\frac{1}{v_A^2} \nabla_{\perp} \Phi \right) + \frac{3}{4} \nabla \nabla_{\perp} \left(\rho_i^2 \frac{1}{v_A^2} \nabla \cdot \nabla_{\perp} \Phi \right) \right] = -\nabla \cdot [\mathbf{b} \nabla^2 (\mathbf{b} \nabla) \Phi]$$

- ▶ The CKA code is used to solve the MHD equations in 3D real magnetic geometry
- ▶ Determines the mode frequency ω and the mode structure $\Phi(r), A_{\parallel}(r)$

$$E_{\parallel} = -\nabla \Phi - \frac{\partial A_{\parallel}}{\partial t} = 0$$

- ▶ Linearised reduced MHD equations
- ▶ Transformed to an eigenvalue problem:

$$\begin{aligned} \omega^2 \left[\nabla \cdot \left(\frac{1}{v_A^2} \nabla_{\perp} \Phi \right) + \frac{3}{4} \nabla \nabla_{\perp} \left(\rho_i^2 \frac{1}{v_A^2} \nabla \cdot \nabla_{\perp} \Phi \right) \right] = & -\nabla \cdot [\mathbf{b} \nabla^2 (\mathbf{b} \nabla) \Phi] \\ & -\nabla \cdot \left[\mathbf{b} \nabla \left(\mu_0 \frac{j_{\parallel}}{B} \mathbf{b} \times \nabla \Phi \right) \right] - \nabla \cdot \left[\frac{\mu_0 \delta p_{\perp}}{B^2} \mathbf{b} \times \nabla B \right] - \nabla \cdot \left[\frac{\mu_0 \delta p_{\parallel}}{B^2} \mathbf{b} \times \kappa \right] \end{aligned}$$

- ▶ The CKA code is used to solve the MHD equations in 3D real magnetic geometry
- ▶ Determines the mode frequency ω and the mode structure $\Phi(r), A_{\parallel}(r)$

$$E_{\parallel} = -\nabla \Phi - \frac{\partial A_{\parallel}}{\partial t} = 0$$

$$\frac{\partial \delta F}{\partial t} + \dot{\mathbf{R}}^{(0)} \frac{\partial \delta F}{\partial \mathbf{R}} + \dot{v}_{\parallel}^{(0)} \frac{\partial \delta F}{\partial v_{\parallel}} = -\dot{\mathbf{R}}^{(1)} \frac{\partial F_0}{\partial \mathbf{R}} - \dot{v}_{\parallel}^{(1)} \frac{\partial \delta F_0}{\partial v_{\parallel}} - \dot{v}_{\perp}^{(1)} \frac{\partial \delta F_0}{\partial v_{\perp}}$$

$$\frac{d\mathbf{R}}{dt} = \frac{\mathbf{B}^*}{B_{\parallel}^*} \left(v_{\parallel} - \frac{q}{m} \langle A_{\parallel} \rangle \right) + \frac{1}{qB_{\parallel}^*} \mathbf{b} \times (\mu \nabla B + q \nabla \langle \Phi - v_{\parallel} A_{\parallel} \rangle)$$

$$\frac{dv_{\parallel}}{dt} = -\frac{1}{m} \frac{\mathbf{B}^*}{B_{\parallel}^*} (\mu \nabla B + q \nabla \langle \Phi - v_{\parallel} A_{\parallel} \rangle),$$

- ▶ The EUTERPE code is used to integrate the equations
- ▶ PIC code, δf method
- ▶ Full particle orbits
- ▶ Currently the linearised version is used
- ▶ Possible to use multiple species

- ▶ The field energy

$$\mathcal{E}_{field} = \frac{1}{2} \int mn_0 \frac{|\nabla_{\perp} \Phi|^2}{B^2} dV + \frac{1}{2\mu_0} \int |\nabla_{\perp} A_{\parallel}|^2 dV$$

- ▶ The growth rate of the mode is given by:

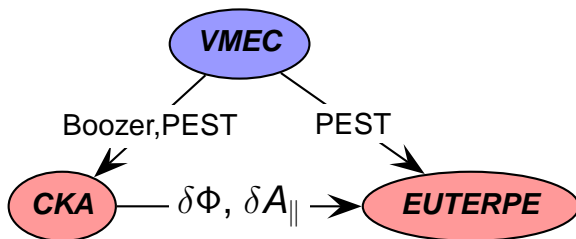
$$\gamma = \frac{1}{2\mathcal{E}_{field}} \frac{\partial \mathcal{E}_{field}}{\partial t} = -\frac{1}{2\mathcal{E}_{field}} \frac{\partial \mathcal{E}_{kin}}{\partial t} = -\frac{1}{2\mathcal{E}_{field}} \int \mathbf{j} \cdot \mathbf{E} d^3r$$

- ▶ Later the pressure perturbation will be calculated

$$\delta p_{\parallel} = \int d\mathbf{v} m v_{\parallel}^2 \delta F, \quad \delta p_{\perp} = \int d\mathbf{v} \mu B \delta F$$

Interface between the codes

- ▶ The magnetic equilibrium is calculated by VMEC
- ▶ The PEST coordinates are implemented in CKA
- ▶ Data transfer routines from CKA to EUTERPE are written
- ▶ The feedback from EUTERPE will be implemented later



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Benchmark with NOVA

GYGLES benchmark

Stellarator results

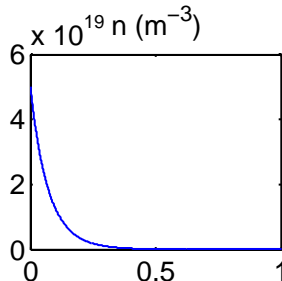
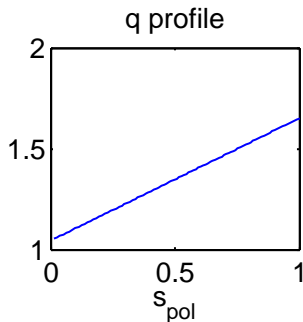
Benchmark 1

Plasma parameters

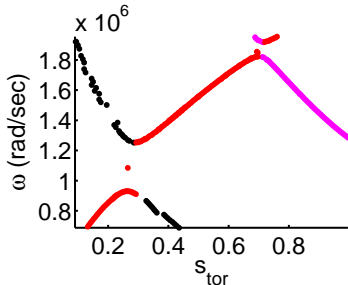
- ▶ Circular tokamak, $R_0 = 4m$, $a=1m$
- ▶ $B_0 = 5T$
- ▶ Deuterium plasma
- ▶ $n_e = 5 \cdot 10^{19}m^{-3}$, constant n profile

Energetic particles

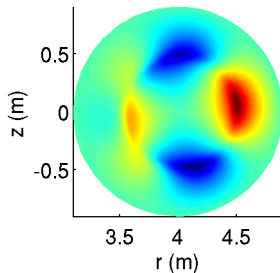
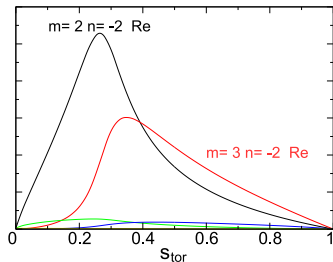
- ▶ Hydrogen
- ▶ $T = 6keV - 1.4MeV$, constant T profile
- ▶ β kept constant
- ▶ density gradient
- ▶ Maxwellian velocity distribution



Mode structure



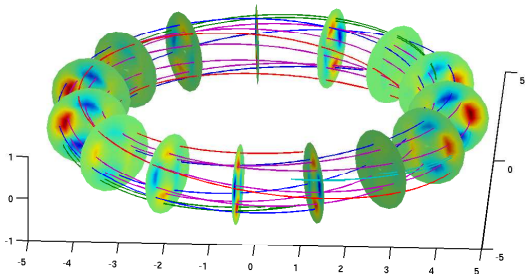
- ▶ Alfvén spectrum calculated by CKA
- ▶ toroidal mode number $n=2$
- ▶ global TAE at the crossing of (2,2) and (2,3) modes



Simulation with EUTERPE

The data is transferred to
EUTERPE

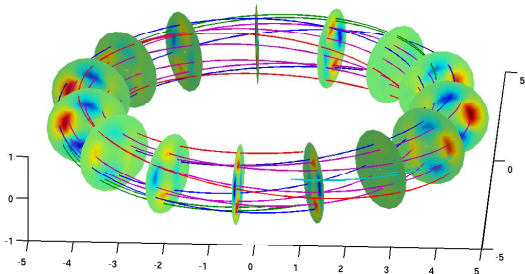
Fast particles started



Simulation with EUTERPE

The data is transferred to
EUTERPE

Fast particles started

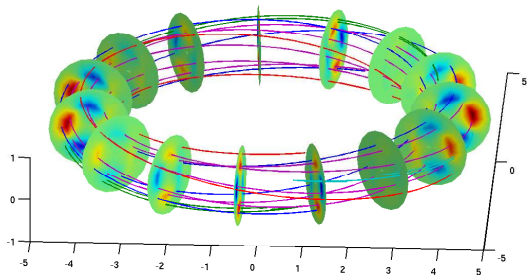


$$\gamma = -\frac{1}{2\mathcal{E}_{field}} \int \mathbf{j} \cdot \mathbf{E} d^3r$$

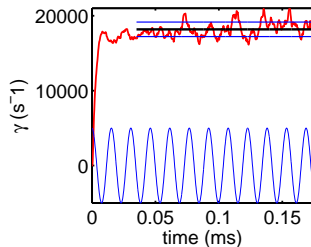
Simulation with EUTERPE

The data is transferred to
EUTERPE

Fast particles started



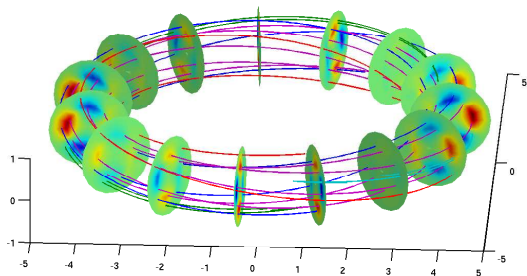
$$\gamma = -\frac{1}{2\mathcal{E}_{field}} \int \mathbf{j} \cdot \mathbf{E} d^3r$$



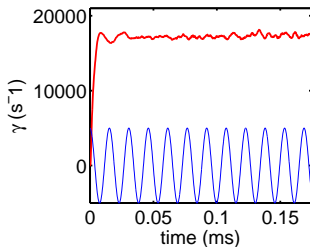
Simulation with EUTERPE

The data is transferred to
EUTERPE

Fast particles started

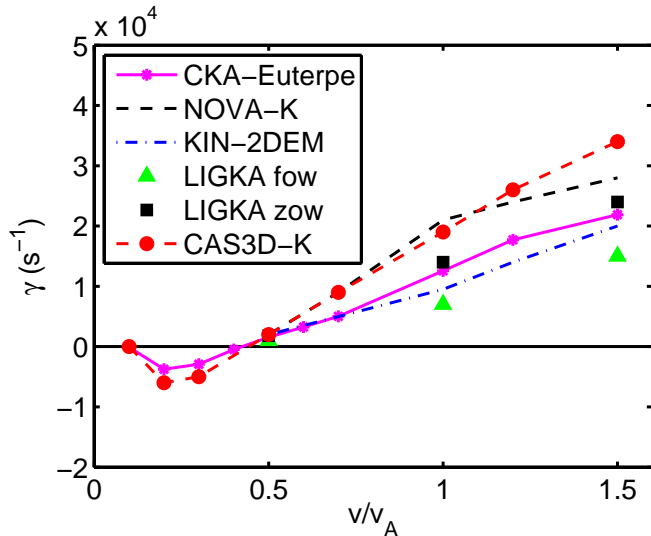


$$\gamma = -\frac{1}{2\mathcal{E}_{field}} \int \mathbf{j} \cdot \mathbf{E} d^3r$$



Stability of the mode

The energetic particle velocity was changed

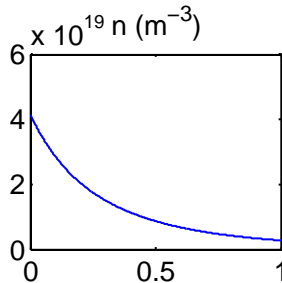
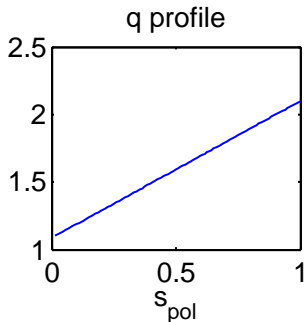


Plasma parameters

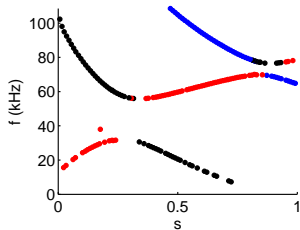
- ▶ Circular tokamak, $R_0 = 3\text{ m}$, $a=1\text{ m}$
- ▶ $B_0 = 1\text{ T}$, $q = 1.1 + \Psi_{pol}$
- ▶ Deuterium plasma
- ▶ $n_e = 4.142 \cdot 10^{19}\text{ m}^{-3}$, constant n profile
- ▶ $T_e = 3.14\text{ keV}$

Energetic particles

- ▶ Deuterium
- ▶ $E_{EP} = 173\text{ keV}$, constant T profile
- ▶ density profile $\sim \exp(-\Psi_{pol}/0.37)$
- ▶ Maxwellian velocity distribution
- ▶ or slowing down $f(v) \sim \frac{1}{v^3 + v_{crit}^3}$

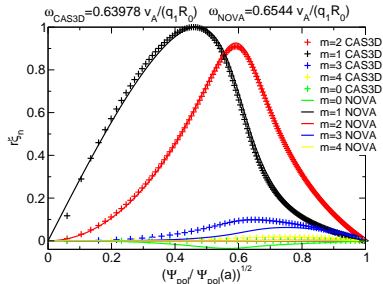
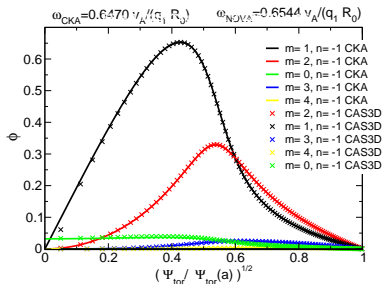


NOVA benchmark, linear n=1 TAE



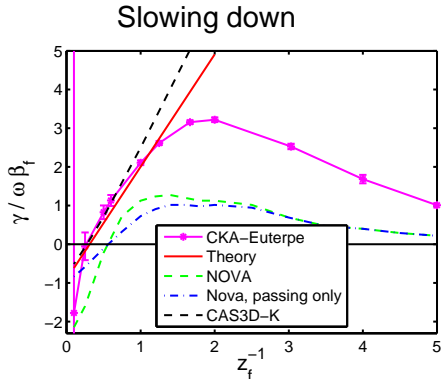
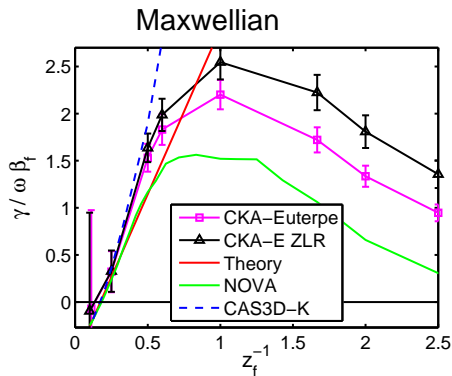
The CKA code is used to calculate the Alfvén mode spectrum

The mode structure from CKA, CAS3D and NOVA:



Stability of the mode

The charge of the fast particles was changed.

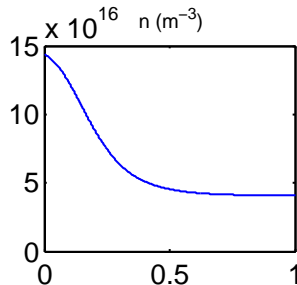
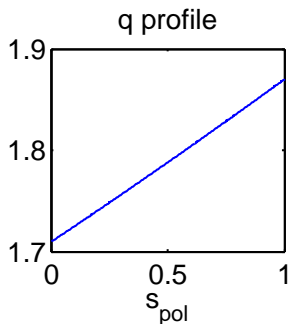


Plasma parameters

- ▶ Circular tokamak, $R_0 = 10m$, $a=1 m$
- ▶ $B_0 = 3T$
- ▶ Hydrogen plasma
- ▶ $n_e = 2 \cdot 10^{19} m^{-3}$, constant n profile
- ▶ $T_e = 1 keV$

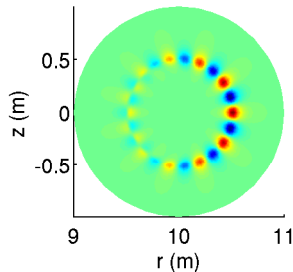
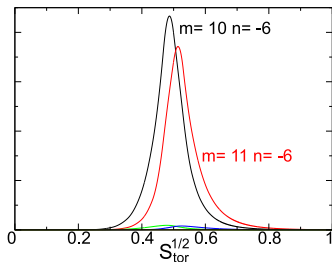
Energetic particles

- ▶ Deuterium
- ▶ $T = 100 keV-800 keV$ constant T profile
- ▶ $n_{0hot} = 1.44 \cdot 10^{17} m^{-3}$, density gradient
- ▶ Maxwellian velocity distribution

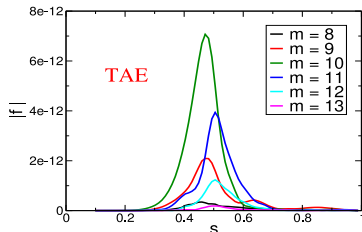


Mode structure, GYGLES benchmark

Mode structure calculated by CKA

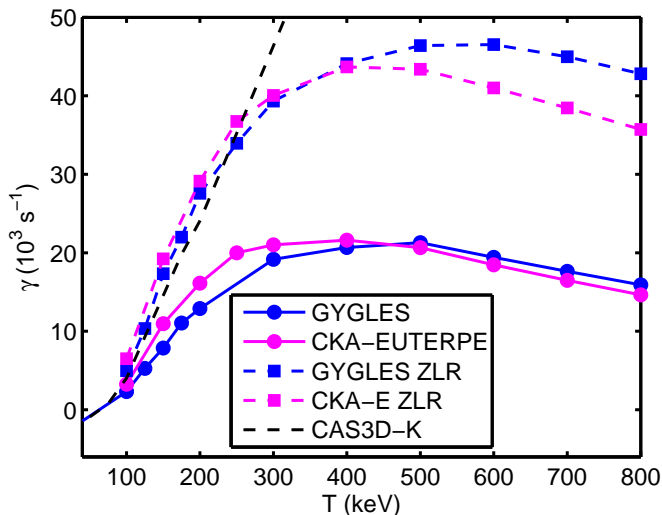


GYGLES mode structure
($T = 400\text{keV}$):



Stability of the mode

The temperature of the fast particles was changed



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Benchmarks

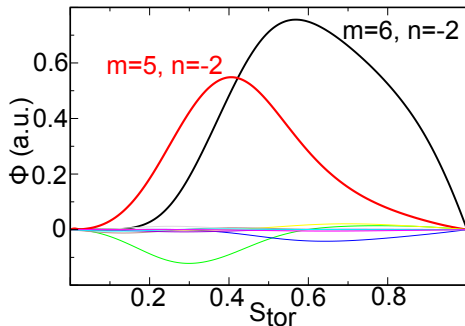
- Benchmark 1

- Benchmark with NOVA

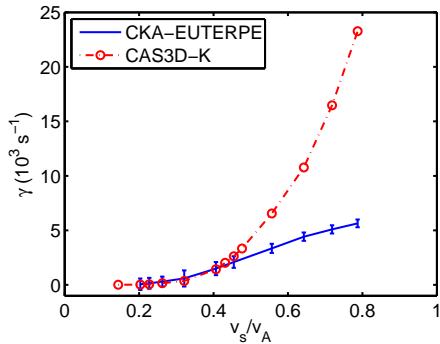
- GYGLES benchmark

Stellarator results

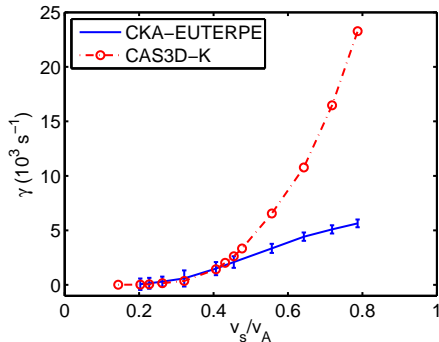
- ▶ Alfvén spectrum determined by CKA
- ▶ Stability of a global TAE is calculated
- ▶ Slowing down distribution function



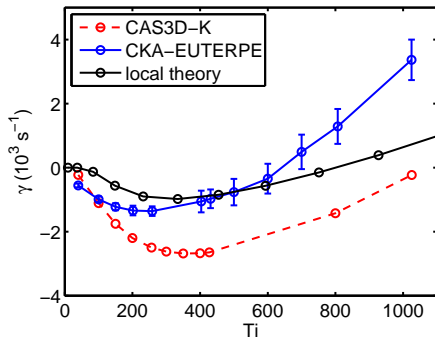
Fast ions



Fast ions



Background ions



CKA-EUTERPE

- ▶ Codes were coupled to create a hybrid code
- ▶ Perturbatively study the stability of Alfvén modes

Linear benchmarks

- ▶ Benchmark 1: successful
- ▶ NOVA benchmark: differences maybe due to the distribution function
- ▶ GYGLES benchmark: successful

Stellarator results

- ▶ W7-AS case calculated
- ▶ FLR and FOW effects are strongly stabilizing